Sunspot observations and the Sunspot Number

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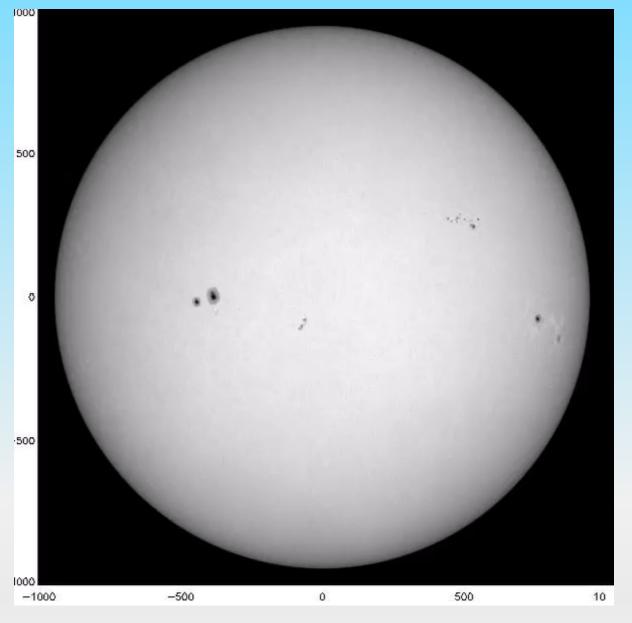
Outline

- Introduction: what are Sunspots?
- Historical overview of Sunspot observations and counting
- Scientific relevance of the Sunspot Number (SSN)
 - In general
 - For climate
- Conclusion
- References for further reading

What are Sunpots?

- Cooler regions of solar photosphere with strong magnetic field
- Typical size: ~ 10⁴ km
- Temperature: ~ 4000 K (instead of 5800 K of the rest of the phostosphere)
- Magnetic field: few Kilogauss
- Sunspots are the most evident sign of the solar magnetic activity

Movie credit: HInode http://solarb.msfc.nasa.gov



- 1610: start of the first systematic observations of Sunspots by Galileo and others
- 1844: discovery of the 11-years cycle by Schwabe
- 1848: the Swiss astronomer Rudolph Wolf (prof. at ETH-Zurich) introduced a Sunspot Number R_i (SSN) (also called Wolf number) based on historical data and its own observations
 - The SSN represents a long and homogenized time series made available to the scientific community for systematic comparisons with several other scientific records (specially meteorological records)

• **Definition** of the international Sunspot Number R_i :

$$R_i = k(10 \times g + s)$$

- g: number of Sunspot Groups
- s: number of individual Sunspots
- k: normalization factor depending on observer

- 1908: Hale observed the magnetic field in Sunspots (through the Zeeman effect)
- 1848 1980: the SSN is determined at ETH Zurich by Wolf and Wolf's successors (Wolfer, Brunner, Waldmeier)
- 1939 & 1957 (IGY): commissioning of the external ETH observing stations in Arosa and Locarno (Specola Solare)
- 1980:
 - The responsibility for the determination of the SSN is transferred to the Royal Observatory in Brussels (in agreement with the International Astronomical Union, IAU)
 - Specola Solare Ticinese in Locarno becomes the world reference station

- July 2015: revision of the SSN
 - Better calibration between different observers
 - Work of several years by many scientists expert on SSN (4 SSN-workshops)
 - Revision lead to a press release by the International Astronomical Union (IAU)
 Assembly in August 2015 in Honolulu

(http://www.iau.org/news/pressreleases/detail/iau1508/)

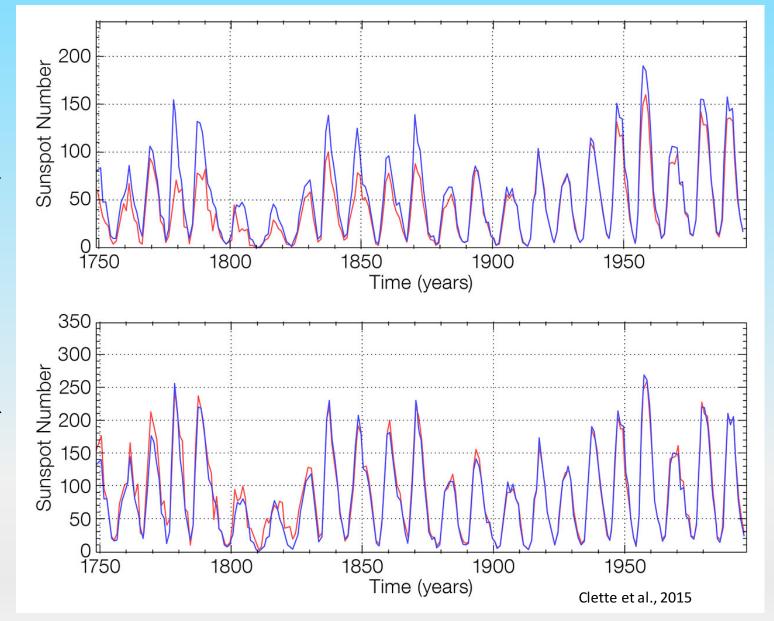
The SSN (1750 – now)

old version

Plot with the new version of the Sunspot Number S_n:

----- SSN

Group number x 19

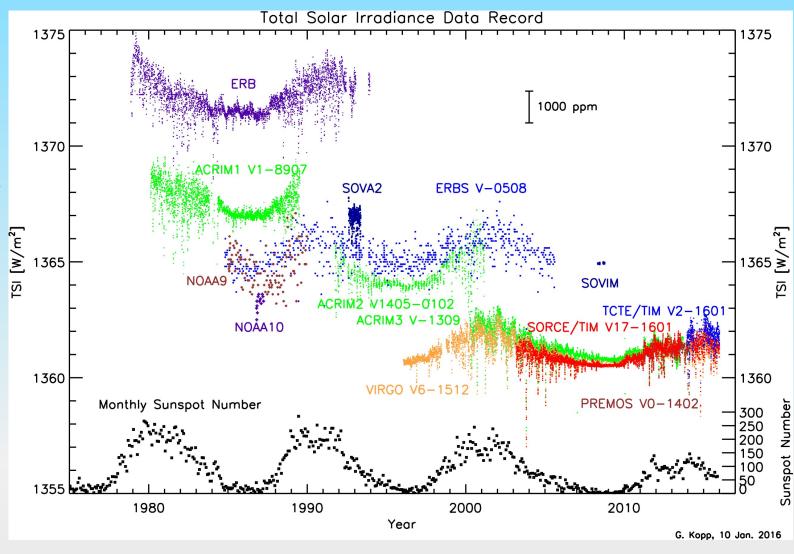


Scientific relevance of the SSN (in general)

- Only direct record of the evolution of the solar cycle over multiple centuries (400 years)
 - Longest scientific experiment still ongoing (B.Owens, Nature, March 2013)
- Multiple applications:
 - Long-term irradiance reconstructions (TSI) (Direct TSI measurements available since 1980!)
 - Calibration of the method based on cosmogenic radioisotope (¹⁴C, ¹⁰Be) production to study the evolution of the solar activity before 1610
 - Earth climate change
 - Radio communication (e.g. HF radio forecast by Swiss Army based on SSN)
 - Infrastructure maintenance (pipelines, electrical power grid)
 - Solar dynamo modelling
- Reference to SSN in > 100 scientific publications / year

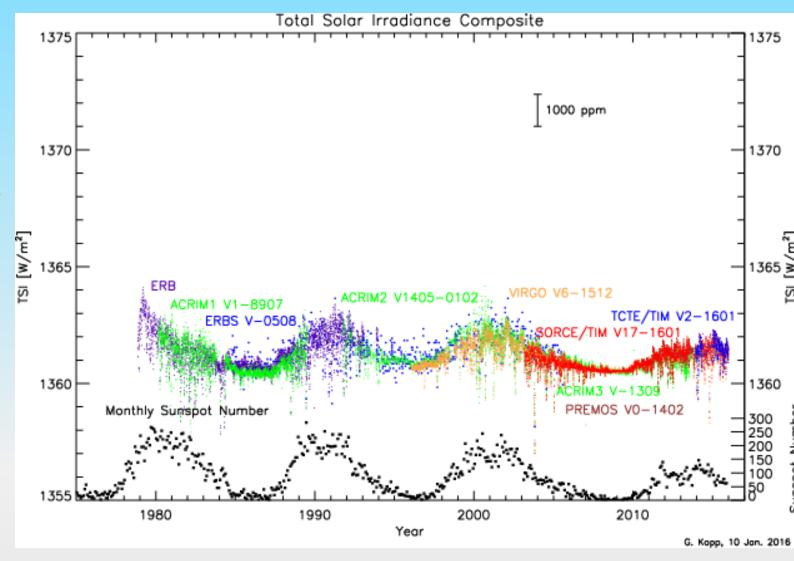
Relation SSN ↔ TSI

- TSI varies of 0.1 % through the solar cycle
- UV- part of TSI more sensitive to solar cycle
- Significant correlation with climate records in the upper atmospheric layers has been reported (e.g. ionosphere, thermosphere)
- Expected change in average Earthsurface temperature due to solar cycle less than 0.1°

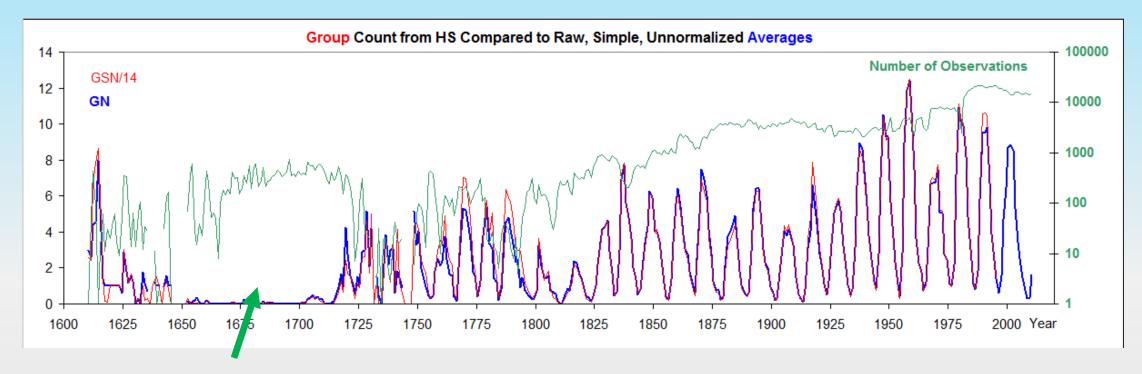


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• Little Ice Age related to Maunder minimum?

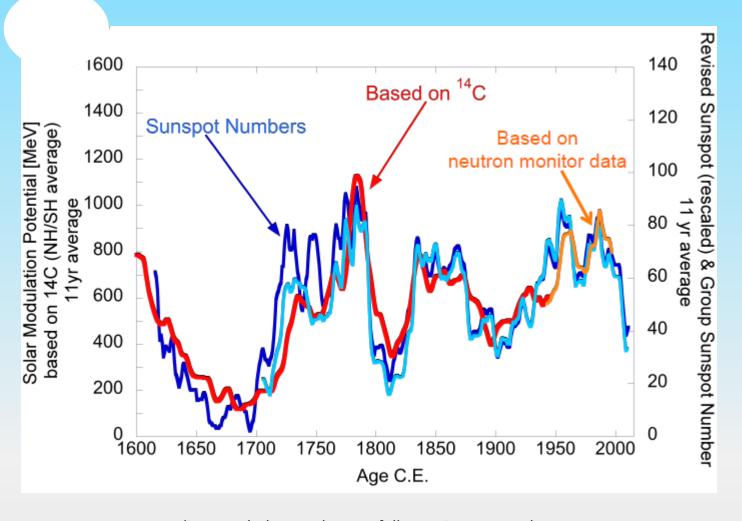


Maunder minimum

Credit: Svalgard & Schatten, Solar Physics, 2015

Cosmogenic nuclei and SSN

- Correlation between ¹⁴C
 atmospheric concentration
 and revised SSN
- Cosmogenic nuclei ¹⁴C and ¹⁰Be allow to study the relation between solar activity and climate before 1610.



Credit: Muscheler et al., AGU fall meeting, December 2015

- Other topics studied:
 - Impact of modulation of cosmic rays through the solar cycle on cloud formation
 - Impact of solar variability on the atmospheric global electric circuit.

• ...

Conclusion

- Due to the relevance of the **Sunspot Number (SSN)** time series for the climate studies over multiple centuries and in particular for the TSI reconstruction, we recommend the inclusion of the SSN as a relevant parameter for the TSI long term studies in the new GCOS implementation plan.
- This recommendation has been proposed in a statement signed by 33 leading scientists in solar physics.

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References for further reading

- Lilensten J. et al., "Earth's climate response to a changing Sun", 2015, ed. EDP-sciences
- Clette F., Svalgaard L., Vaquero J.M., & Cliver E.W., "Revisiting the Sunspot Number. A 400-Year Perspective on the Solar Cycle", 2014, Space Science Review, 186, 35
- Cliver E.W., Clette F., Svalgaard L. & Vaquero J.M., "Recalibrating the Sunspot Number (SN): The 3rd and 4th SN Workshops", 2015, Central European Astrophysical Bulletin, 39, 1
- Ermolli I., Shibasaki K., Tlatov A., & van Driel-Gesztelyi L., "Solar Cycle Indices from the Photosphere to the Corona: Measurements and Underlying Physics", 2014, Space Science Review, 186, 105
- Cortesi S., Cagnotti M., Bianda M., Ramelli R., Manna A., "Sunspot Observations and Counting at Specola Solare Ticinese in Locarno since 1957", 2016, in press on Solar Physics
- Solar Physics (journal), Topical issue about Sunspot Number (2016 in press)
- http://www.iau.org/news/pressreleases/detail/iau1508/